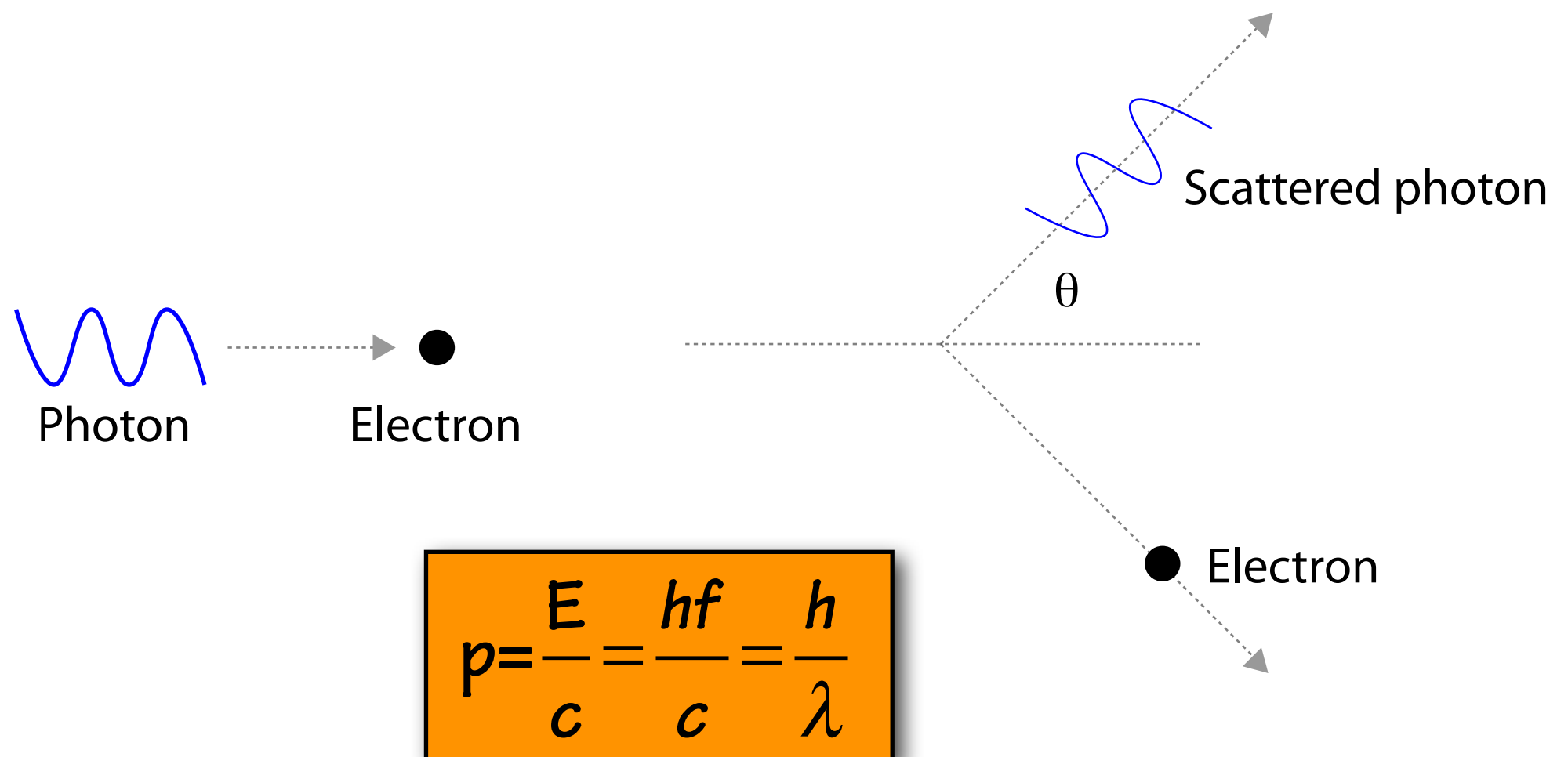


Matter & waves

- Photon momentum
- Photon momentum calculations
- De-Broglie wavelength
- Electron diffraction
- Emission & absorption spectra
- Energy levels Bohr's model of electrons
- Hydrogen emission spectrum
- Mercury emission spectrum
- Electron waves

Photon momentum

- The Compton effect: x-ray photons will cause a recoil of electrons.
- Photons carry momentum.
- The gain in momentum of the electron causes a loss of momentum & energy of the photon: the scattered photon has a longer wavelength.



Photon momentum calculations

- eg. for violet light $\lambda = 4 \times 10^{-7} \text{ m}$
- Recall that a 1 mW violet laser emits $\sim 2 \times 10^{15}$ photons per second:

Momentum per photon:

$$p = \frac{h}{\lambda}$$

$$p = \frac{6.63 \times 10^{-34} \text{ Js}}{400 \times 10^{-9} \text{ m}}$$

$$p = 1.7 \times 10^{-27} \text{ kgm/s}$$

Force when photons are absorbed
over a second

$$F = \frac{\Delta p}{\Delta t}$$

$$F = (1.7 \times 10^{-27} \text{ Ns / photon}) \\ \times (2 \times 10^{15} \text{ Photons / s})$$

$$F = 3 \times 10^{-12} \text{ N}$$



Solar sail

De-Broglie wavelength

- Double slit experiment with electrons showed interference patterns.
- Evidence of electrons showed wave properties.
- Louis De-Broglie's idea: every particle has a De-Broglie wavelength dependent upon its momentum
- larger p = small λ
- eg. a tennis ball $p = (0.05 \text{ kg})(30 \text{ m/s}) = 1.5 \text{ kgm/s}$

$$p = \frac{h}{\lambda} \longrightarrow \lambda = \frac{h}{p}$$

$$\lambda = \frac{6.63 \times 10^{-34} \text{ Js}}{1.5 \text{ kgm/s}}$$

$$\lambda = 4.0 \times 10^{-34} \text{ m}$$

Too small a wavelength to be noticeable:
the ball acts as a particle

Electron diffraction

- Electrons carry much less momentum & have more noticeable wave properties.
- $E = 100 \text{ eV} (= 1.6 \times 10^{-17} \text{ J})$, $m \sim 9.1 \times 10^{-31} \text{ kg}$
- This is of a size comparable to atoms, so diffraction from crystal structures is possible.

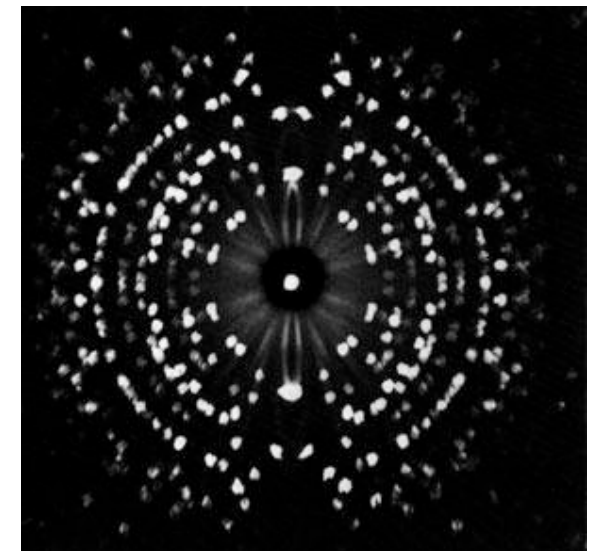
$$E_k = \frac{1}{2}mv^2 \rightarrow v = \sqrt{\frac{2E_k}{m}} \quad v = \sqrt{\frac{2 \times 1.6 \times 10^{-17} \text{ J}}{9.1 \times 10^{-31} \text{ kg}}} = 5.9 \times 10^6 \text{ m/s}$$

$$p = mv = (9.1 \times 10^{-31} \text{ kg}) \times (5.9 \times 10^6 \text{ m/s})$$

$$p = mv = 5.4 \times 10^{-24} \text{ kgm/s}$$

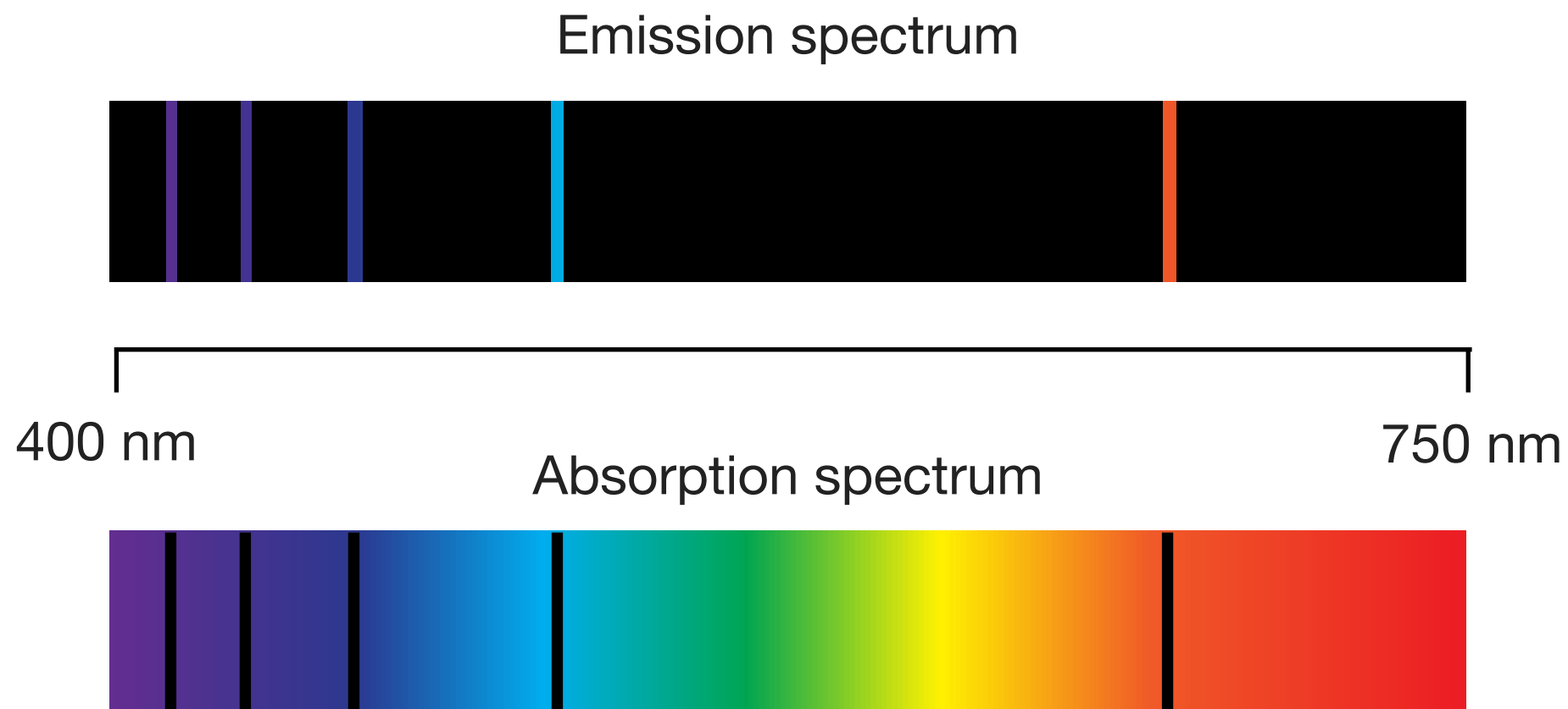
$$\lambda = \frac{6.6 \times 10^{-34} \text{ Js}}{5.4 \times 10^{-24} \text{ kgm/s}} = 1.2 \times 10^{-10} \text{ m}$$

$\sim 10^{-10} \text{ m}$ is considered the boundary
of wave or particle behaviour



Emission & absorption spectra

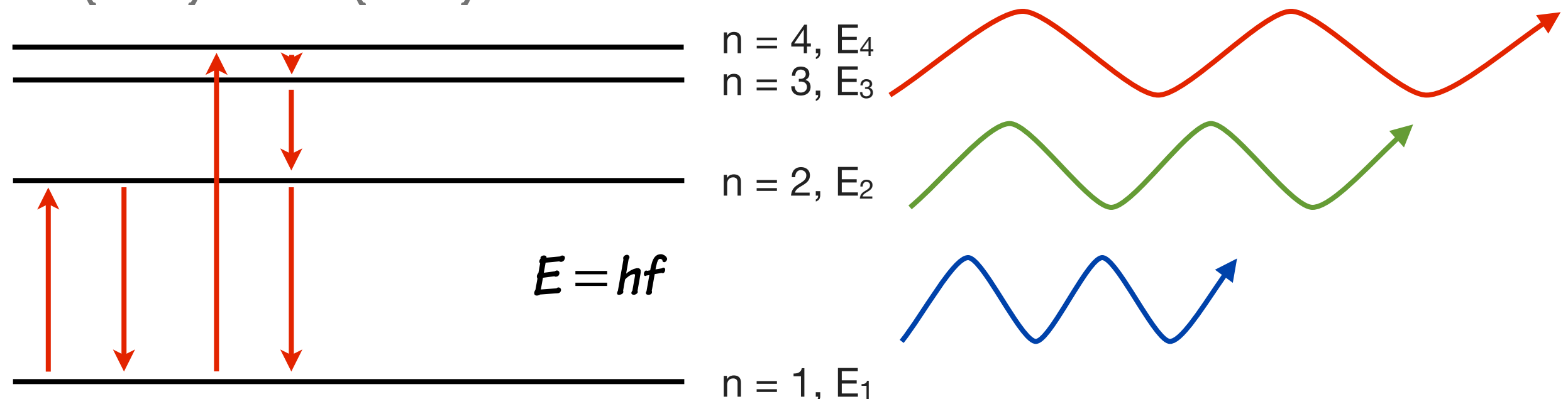
- Each element has a unique “fingerprint” - the spectrum of light that is produced when the gas is heated.
- Only certain wavelengths of light are emitted.
- If the light passes through the gas, those same wavelengths will be absorbed.



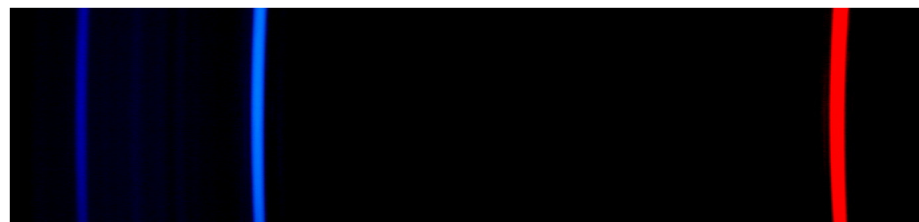
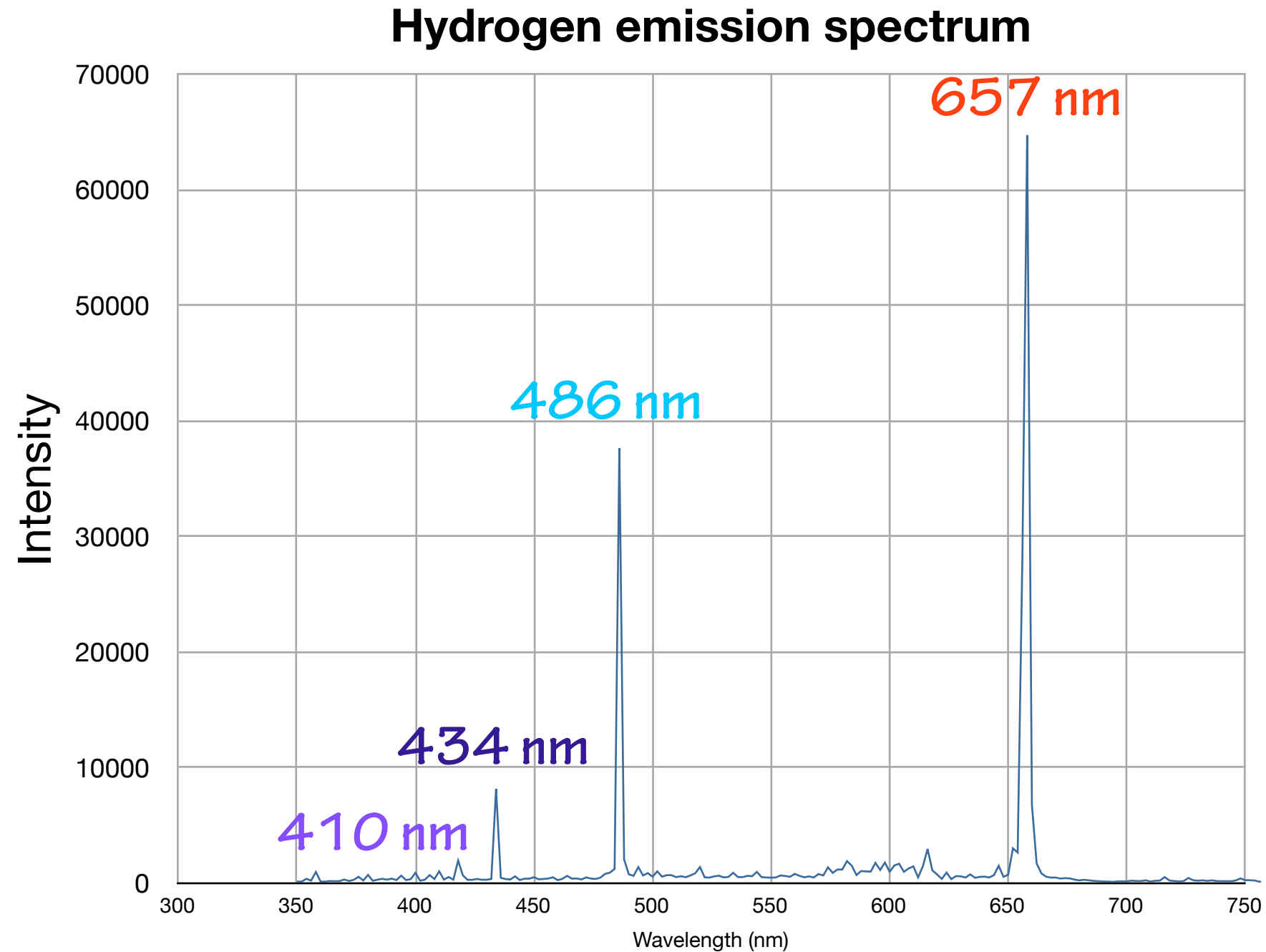
The wavelengths emitted are the same that are absorbed by the element.

Energy Levels - Bohr's model of electrons

- The "ground state" (E_1) is the lowest energy state.
- If an electron absorbs enough energy from a photon, then it can move up into a higher energy state.
- Likewise, an electron may emit a photon and drop down to a lower energy state. The photon energy is the difference of electron energies.
- This explains "Spectral lines": discrete wavelengths emitted (or absorbed) by a particular atom. (Used to identify what elements are present!)
- The "ionisation energy" is the energy required to bring an electron from E_1 ($n=1$) to E_∞ ($n=\infty$).

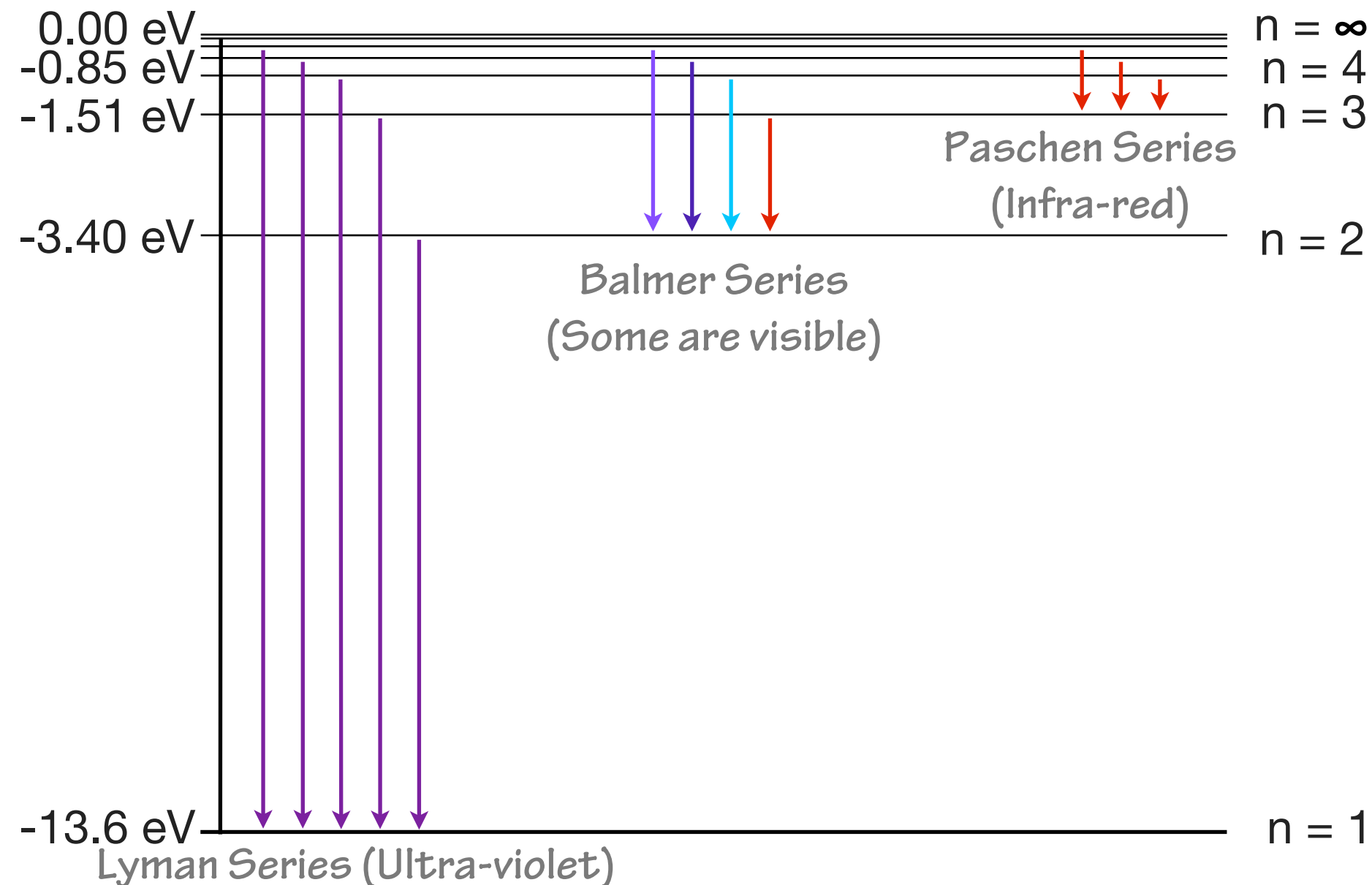


Hydrogen emission spectrum



Hydrogen emission series

- Distinct spectral lines are visible from hydrogen that correspond to the emission of energy from excited states to lower energy states.
- Each series represents the emission to one particular state.



$$E = \frac{-13.6 \text{ eV}}{n^2}$$

n	E
∞	0 eV
10	-0.14 eV
9	-0.17 eV
8	-0.21 eV
7	-0.28 eV
6	-0.38 eV
5	-0.54 eV
4	-0.85 eV
3	-1.51 eV
2	-3.40 eV
1	-13.60 eV

Hydrogen emission series - The Balmer Series

- The visible emission lines of Hydrogen are from the Balmer series, as electrons make the transition to the $n = 2$.

$$3 \rightarrow 2 \quad E = -3.40 \text{ eV} - (-1.51 \text{ eV}) = -1.89 \text{ eV}$$

$$\lambda = \frac{hc}{E} = \frac{(4.14 \times 10^{-15} \text{ eVs})(3.00 \times 10^8 \text{ m/s})}{1.89 \text{ eV}}$$

$$= 657 \text{ nm (red)}$$

$$4 \rightarrow 2 \quad E = -3.40 \text{ eV} - (-0.85 \text{ eV}) = -2.55 \text{ eV}$$

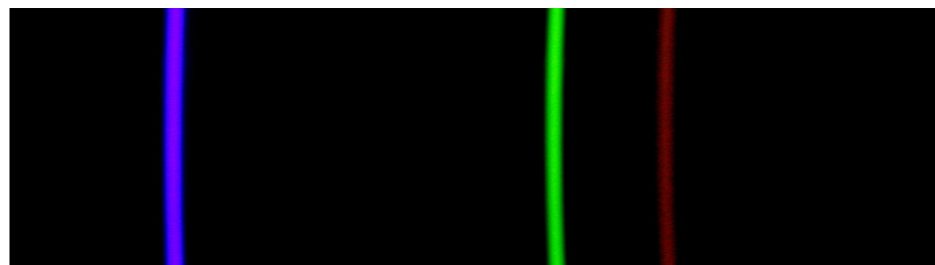
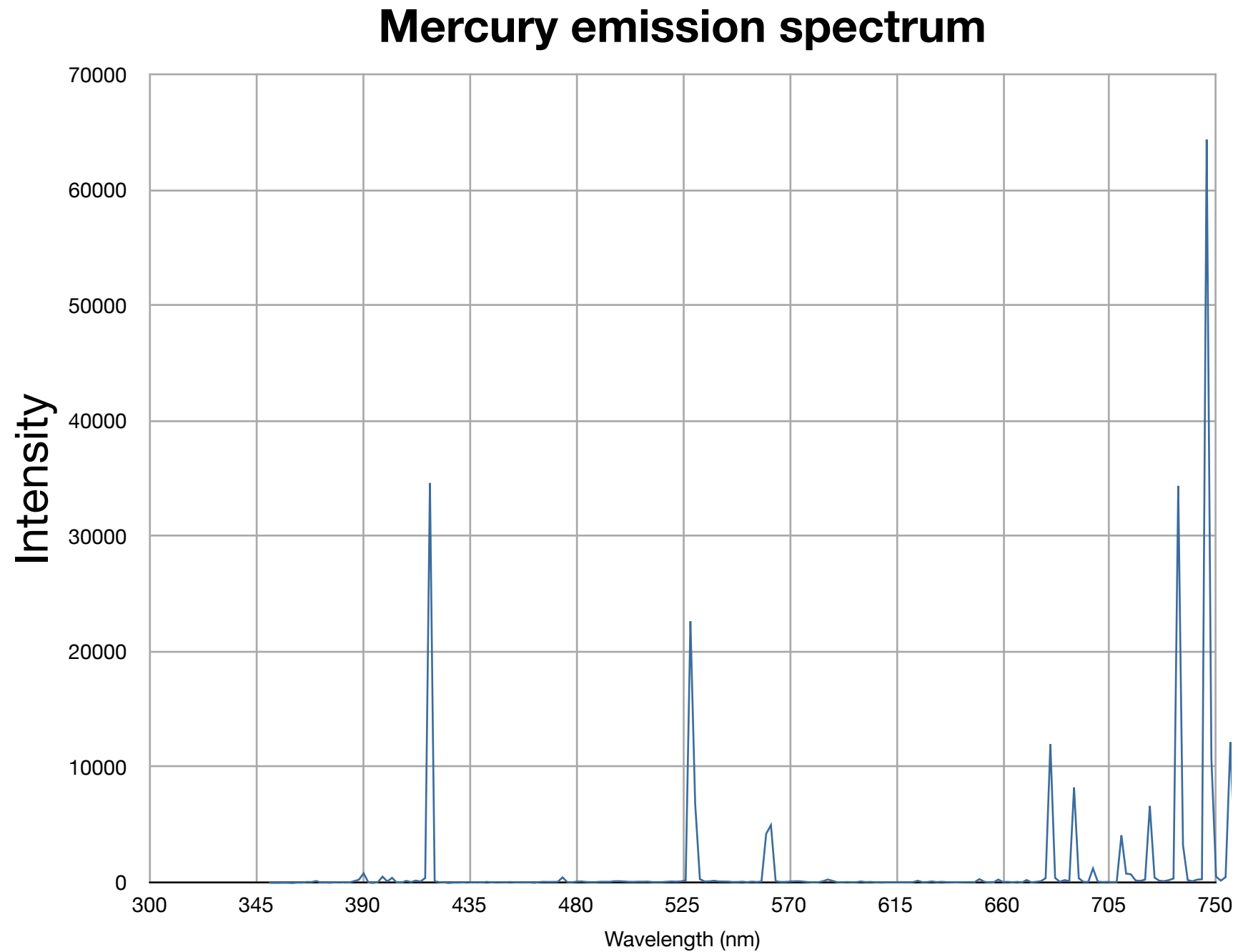
$$\lambda = \frac{hc}{E} = \frac{(4.14 \times 10^{-15} \text{ eVs})(3.00 \times 10^8 \text{ m/s})}{2.55 \text{ eV}}$$

$$= 486 \text{ nm (cyan)}$$

The visible spectrum ranges from 400 nm (3.1 eV)
to 750 nm (1.7 eV)

n	E
∞	0 eV
10	-0.14 eV
9	-0.17 eV
8	-0.21 eV
7	-0.28 eV
6	-0.38 eV
5	-0.54 eV
4	-0.85 eV
3	-1.51 eV
2	-3.40 eV
1	-13.60 eV

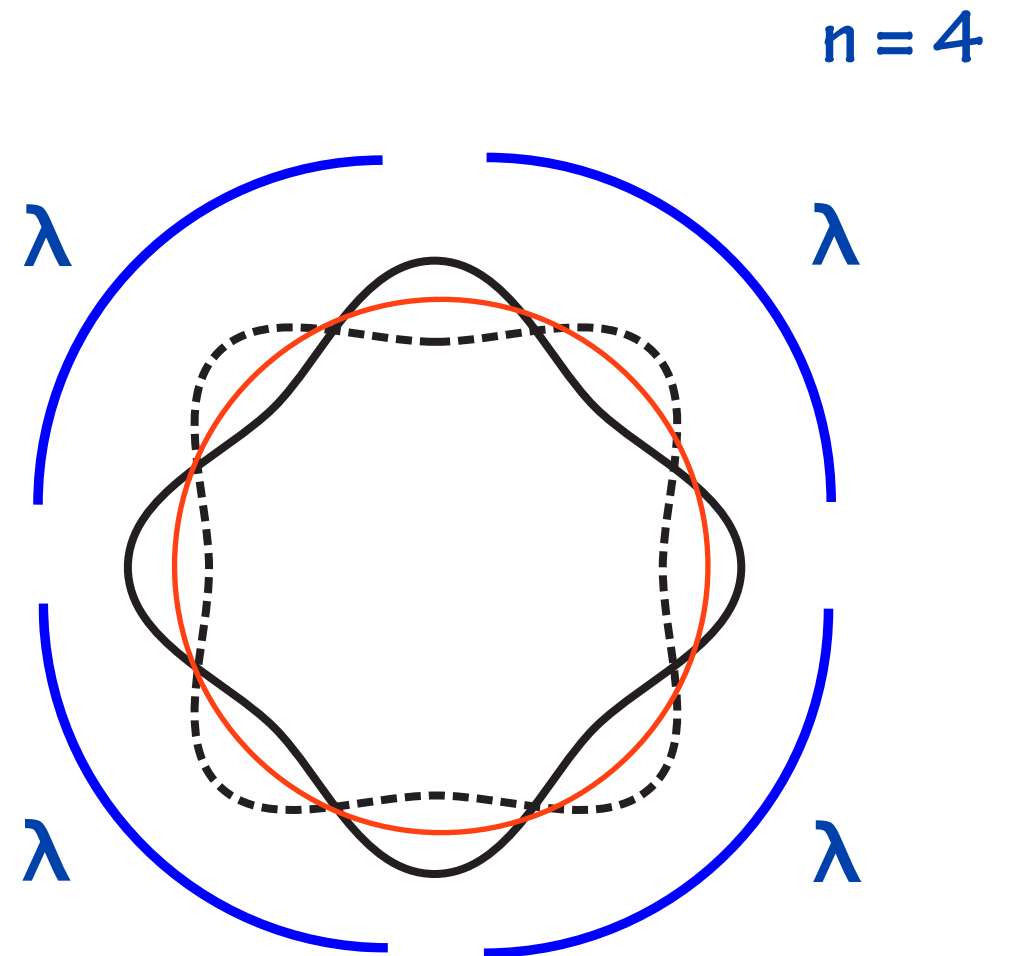
Mercury emission spectrum



Electron waves

- Bohr developed a model of the atom in which the electrons could only occupy orbits of certain radii that correspond to discrete energy levels.
- De Broglie's hypothesis: the electron can only exist in an orbit if it forms a standing wave.
- The radius must be a multiple of a wavelength.
- If the radii are quantised, then so is the energy of the electrons.
- This explains why only certain energy levels & emission spectra are possible.

$$2\pi r = n\lambda = \frac{nh}{p}$$



Bohr's theory

Light - a wave or a particle?

- Light clearly shows wave properties (interference) & particle properties (photoelectric effect).
- Likewise, matter shows wave properties (electron diffraction).
- So it seems that we need to use both models to properly describe light.
- But as a general rule, the larger an object is - the more particle like it will be acting.
- For electromagnetic radiation - the shorter the wavelengths will act more like particles.
- Remember that waves or particles are just models to try to describe a physical situation.